



## Supplementary Materials for

### **Efficient and stable large-area perovskite solar cells with inorganic charge extraction layers**

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**Correction:** Reference 34 was mistakenly listed as X. F. Chen *et al.*, *Int. J. Clin. Exp. Med.* **8**, 5044–5054 (2015). The correct reference 34 is W. Chen *et al.*, *Energy Environ. Sci.* **8**, 629–640 (2015). The PDF has been corrected.

## Materials and Methods

### Materials

All the chemicals were used as received, including  $\text{PbI}_2$  (99%, Sigma–Aldrich),  $\text{CH}_3\text{NH}_3\text{I}$  (> 98%, Tokyo Chemical Industry Co., Japan), Titanium isopropoxide (99.999%, Sigma–Aldrich), niobium ethoxide (99.95%, Sigma–Aldrich), nickel acetylacetonate (95%, Sigma–Aldrich),  $\text{PC}_{61}\text{BM}$  (99.5%, Lumtec Co., Taiwan). Magnesium acetatetetrahydrate (99%), lithium acetate (99%) and super dehydrated solvents of dimethylsulfoxide (DMSO), toluene, chlorobenzol, acetonitrile, methanol and ethanol, were all purchased from Wako Co., Japan.

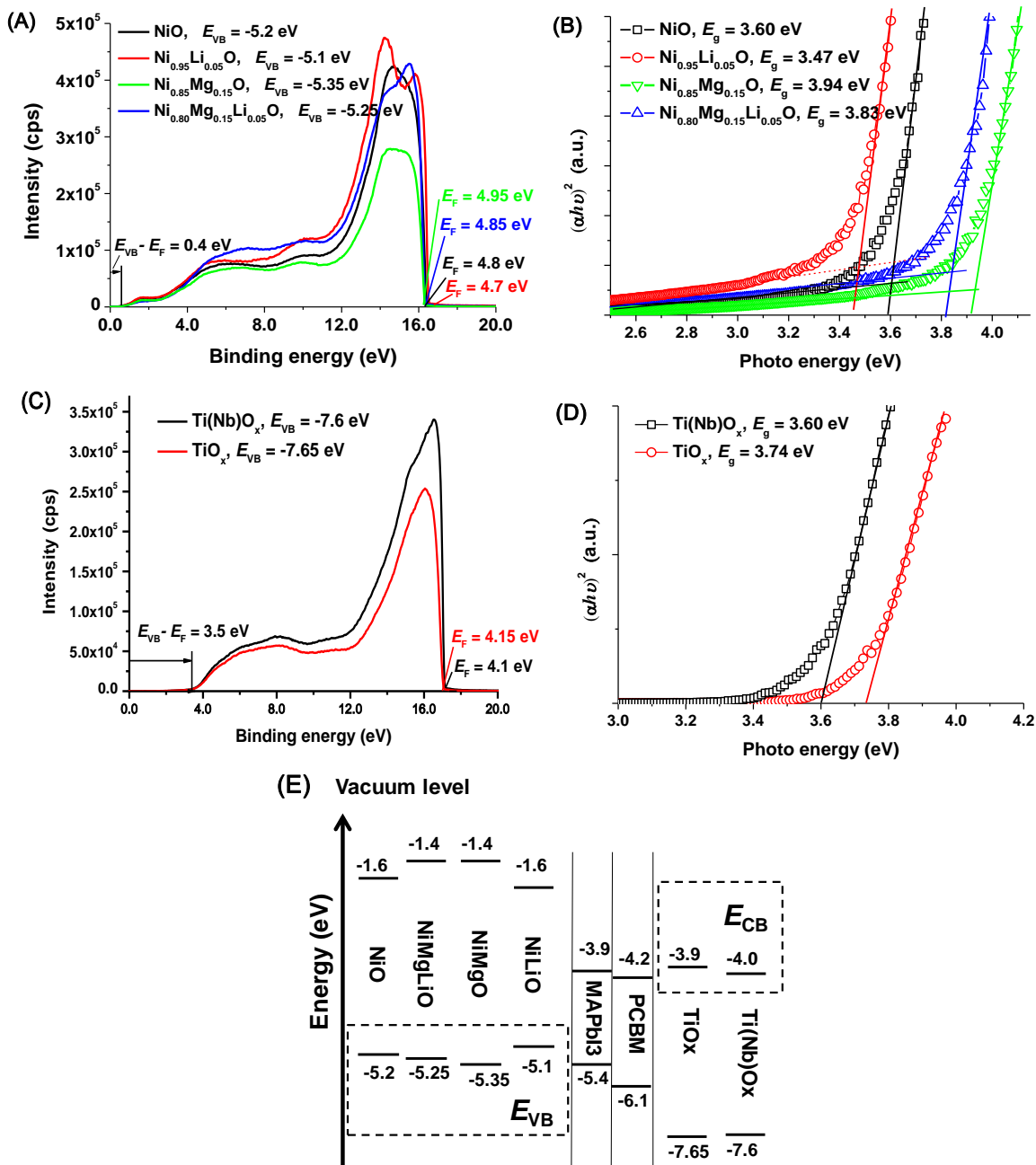
### Thin Film and Solar Cell Fabrication

FTO glasses (Nippon Sheet Glass Co., Japan) with a sheet resistance of  $8\text{--}10\ \text{ohm}^{-2}$  and an optical transmission of greater than 80% in the visible range were used. After thoroughly washing with cleaning fluid, deionized water, ethanol and acetone, the clean FTO glasses were placed on a hotplate at the temperature of  $500\ ^\circ\text{C}$ . 30 ml of an acetonitrile/ethanol (with 95: 5 volume ratio) solution of nickel acetylacetonate (or with 15 mol% magnesium acetatetetrahydrate and 5 mol% lithium acetate) was sprayed within 10 min by an air nozzle (with 0.2 mm caliber) onto the hot FTO glasses at a distance of about 20 cm above. The total metal ion concentration was set at  $0.02\ \text{mol L}^{-1}$ . After spraying, the film was further treated at  $500\ ^\circ\text{C}$  for another 20 min to promote NiO crystallization. NiO film thickness could be tuned from 10 to 40 nm by varying the solution volume from 15 to 60 ml. After cooling, the NiO coated FTO glasses were transferred quickly to the  $\text{N}_2$  filled glove box for  $\text{MAPbI}_3$  perovskite layer deposition. The method for perovskite layer deposition was modified slightly from previous report(6): first, a  $80\ \mu\text{l}$  1.45 M DMSO solution of  $\text{PbI}_2/\text{MAI}$  mixture was spread onto  $70\ ^\circ\text{C}$  FTO glass ( $2.4\times 2.4\ \text{cm}^2$  in size); then, the spin-coater was started at a rotation speed of 1000 rpm for 10 seconds and 5000 rpm for another 30 seconds.  $800\ \mu\text{l}$  toluene was drop-casted quickly 10 seconds after the 5000 rpm spin-coating started. The perovskite films were then heated at  $100\ ^\circ\text{C}$  on a hotplate for 10 min. After cooling, a chlorobenzol solution of PCBM ( $20\ \text{mg ml}^{-1}$ ) was spin-coated on top of the perovskite film at the rotation speed of 1000 rpm for 30 seconds. Subsequently,  $100\ \mu\text{l}$   $\text{TiO}_x$  precursor solution, containing titanium isopropoxide (or with 5 mol% niobium ethoxide) diluted in methanol with volume ratio of 1/200–1/600, was drop-casted slowly onto the surface of PCBM layer at the rotation speed of 6000 rpm. After the film was dried at  $70\ ^\circ\text{C}$  for 10 minutes,  $150\ \mu\text{l}$   $\text{H}_2\text{O}$ /methanol mixed solution (with 1: 99 volume ratio) was drop-casted slowly at the rotation speed of 6000 rpm for 30 seconds, to promote hydrolysis of previously deposited titanium isopropoxide. The film was then dried at  $70\ ^\circ\text{C}$  for 30 min to promote condensation of titanium hydroxide, leading to  $\text{TiO}_x$  formation. At last, one batch of films were transferred to the evaporator chamber, 100 nm thick Ag contacts were deposited under high vacuum ( $< 3\times 10^{-4}\ \text{Pa}$ ). For solar cell sealing, a cavity glass with UV glue on the edges was sandwiched with the front FTO glass with the active films in-between. After exposure to 300 W UV light for 15 seconds, the sealing process was terminated. All of these processes were done in glove box.

## Characterizations

SEM images were obtained by using a JSM-6500F field-emission scanning electron microscope. TEM observations were carried out on a JEM-ARM200F transmission electron microscope. For cross sectional observation of NiO films, the samples were prepared by using a FIB cutting technique. Electric conductivity of the charge extraction films (NiO based films on FTO glass, and TiO<sub>x</sub> based films on PCBM covered ITO glass) was investigated by using the contact-current mode of a scanning probe microscope (SPM; JEOL JSPM-5200). The XPS and UPS measurements were performed in a Kratos Ultra Spectrometer (AXIS-ULTRA DLD-600W) using monochromatized Al K $\alpha$  X-ray photons ( $h\nu = 1486.6$  eV for XPS) and a HeI (21.2 eV for UPS) discharge lamp. A Ecopia HMS 5500 Hall system was applied to measure the room-temperature conductivity and Hall coefficient ( $RH$ ) by using the van der Pauw method with a magnetic field strength of 0.550 T, and the carrier concentration ( $n$ ) was calculated according to the equations:  $|RH| = 1/ne$ , where  $e$  is the electrical charge of the electron. The UV-Vis spectra were measured on a Shimadzu UV/Vis 3600 spectrophotometer with an integrating sphere. The thickness of NiO based films was observed and measured by SEM. The thickness of TiO<sub>x</sub> based films (on top of PCBM covered ITO glasses) was measured by a Surfcom 1400A surface profiler (Tokyo Seimitsu Co. Ltd.) deducting the thickness of PCBM underlayers.

The current-voltage characteristics were measured using a black metal mask with an aperture area of 0.09 or 1.02 cm<sup>2</sup> under standard air mass 1.5 sunlight (100 mW cm<sup>-2</sup>, WXS-155S-10: Wacom Denso Co., Japan). Monochromatic incident photon-to-current conversion efficiency (IPCE) spectra were measured with a monochromatic incident light of  $1 \times 10^{16}$  photons cm<sup>-2</sup> in director current mode (CEP-2000BX, Bunko-Keiki). The light soaking stability was tested in a solar cell light resistance test system (Model BIR-50, Bunkoh-Keiki Co., LTD) equipped with a Class AAA solar simulator; <420 nm UV light was cut off with an optical filter. The time-resolved photoluminescence characterizations were done on an Edinburg FLS 920 (Edinburg90 Co. LTD), and the excitation was provided by a picosecond pulsed light emitting diode (EPLED-445) at the wavelength of 445 nm with pulse width of 95.3 ps. Transient photovoltage/photocurrent decay measurements were done on a home-made system. A white light bias on the sample was generated from an array of diodes. Red light pulse diodes (0.05 s square pulse width, 100 ns rise and fall time) controlled by a fast solid-state switch were used as the perturbation source. The transient photocurrent was measured using 20 ohm external series resistance to operate the device in short circuit. Similarly, transient photovoltage was measured using 1 M $\Omega$  external series resistance to operate the device in open circuit. The voltage dynamics on the resistors were recorded on a digital oscilloscope (Tektronix MDO3032). The perturbation red light source was set to a suitably low level to the white diodes array with light intensity equivalent to 100 mW cm<sup>-2</sup> of a standard solar simulator.



**Fig. S1.**

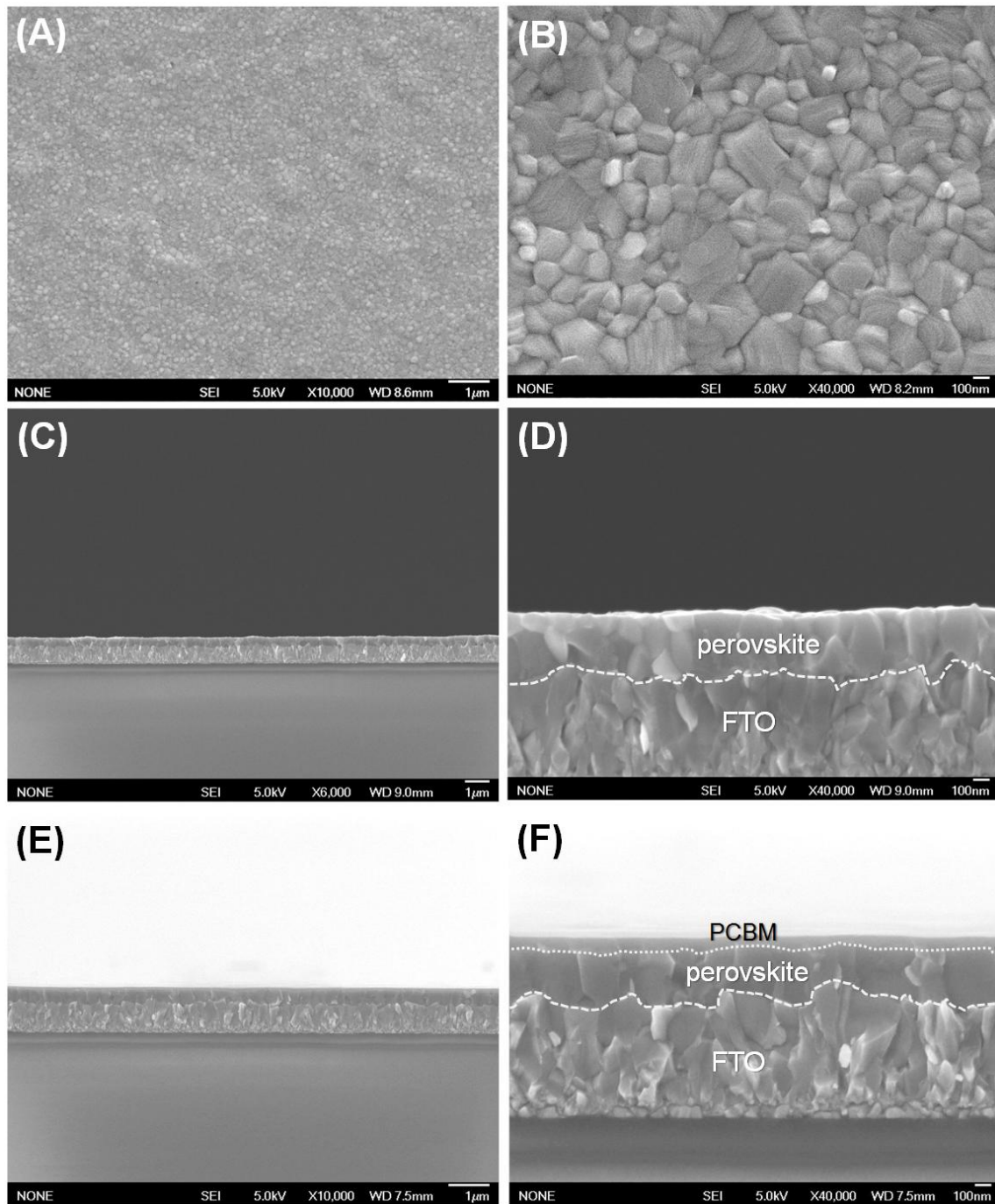
(A) UPS and (B) UV-Vis spectra of the NiO based charge extraction layers with and without doping.  $NiMgLiO$ ,  $NiMgO$  and  $NiLiO$  represent  $Mg_{0.15}Li_{0.05}Ni_{0.80}O$ ,  $Mg_{0.15}Ni_{0.85}O$  and  $Li_{0.05}Ni_{0.95}O$ , respectively. (C) UPS and (D) UV-Vis spectra of the  $TiO_x$  based charge extraction layers with and without doping. (E) Summary on the calculated band edge positions of the charge extraction layers with respect to MAPbI<sub>3</sub> and PCBM.

The work function, or Fermi level ( $E_F$ ) of the charge carrier extraction layers are obtained by subtracting the binding energies of the secondary electron cutoffs (in the range of 16–18 eV) from the excitation energy (21.2 eV) of HeI UPS spectra. The energy

difference between the valance band maximum ( $E_{VB}$ ) and  $E_F$  was derived from the low binding energy tails (in the range below 4 eV). Conduction band ( $E_{CB}$ ) edge values were obtained by adding the optical band gaps ( $E_g$ ) to  $E_{VB}$ .

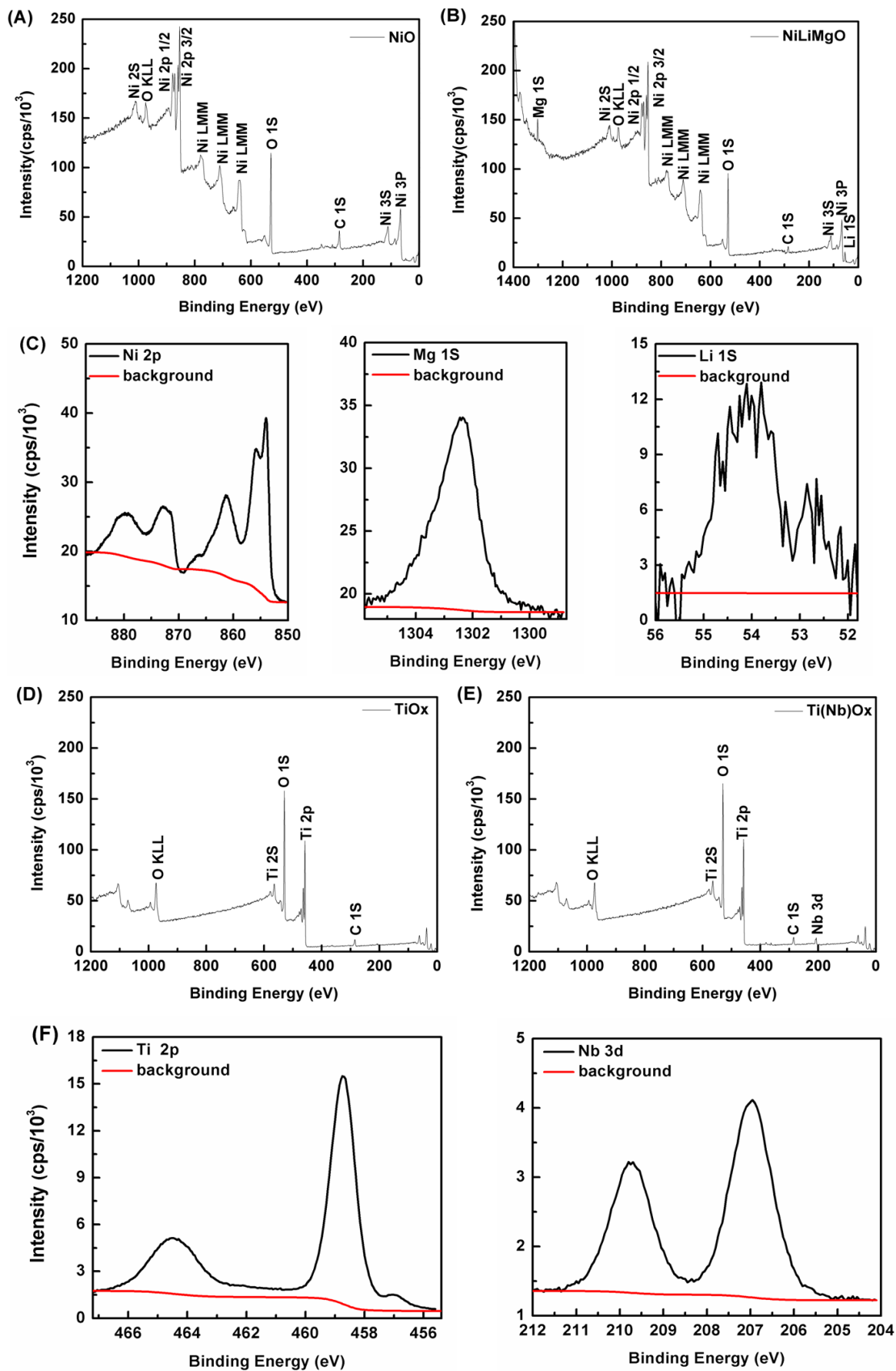
Both NiO and MgO exhibit rock-salt structures with a small (0.8%) lattice mismatch. The lattice constant for NiO and MgO is 4.177 Å and 4.213 Å, respectively. Therefore, solid solution of  $Ni_xMg_{1-x}O$  ( $x = 0-1$ ) could be obtained as reported in the literatures. At the same time, according to Vegard's law, the band gaps of the  $Ni_xMg_{1-x}O$  alloys could be continuously tuned by varying “x”(38). The value of  $Ni_{0.85}Mg_{0.15}O$  determined from this work with a wider band gap and about 0.15 eV deeper valance band are consistent with the literature (37).

$Ni_{1-x}Li_xO$  also has very high doping tolerance. The NiLiO alloy with “x” value up to 0.25 has been reported without crystal structure change (48). Suitable Li doping could greatly enhance the conductivity of NiO, allowing  $Ni_{1-x}Li_xO$  to be a candidate for a p-type TCO material. The optimal “x” value for the highest conductivity as reported in the literatures is around 3–10% (49). In this work,  $x = 5\%$  was selected as the target composition, in order to prevent too much change on  $E_{VB}$  and  $E_g$ .



**Fig. S2**

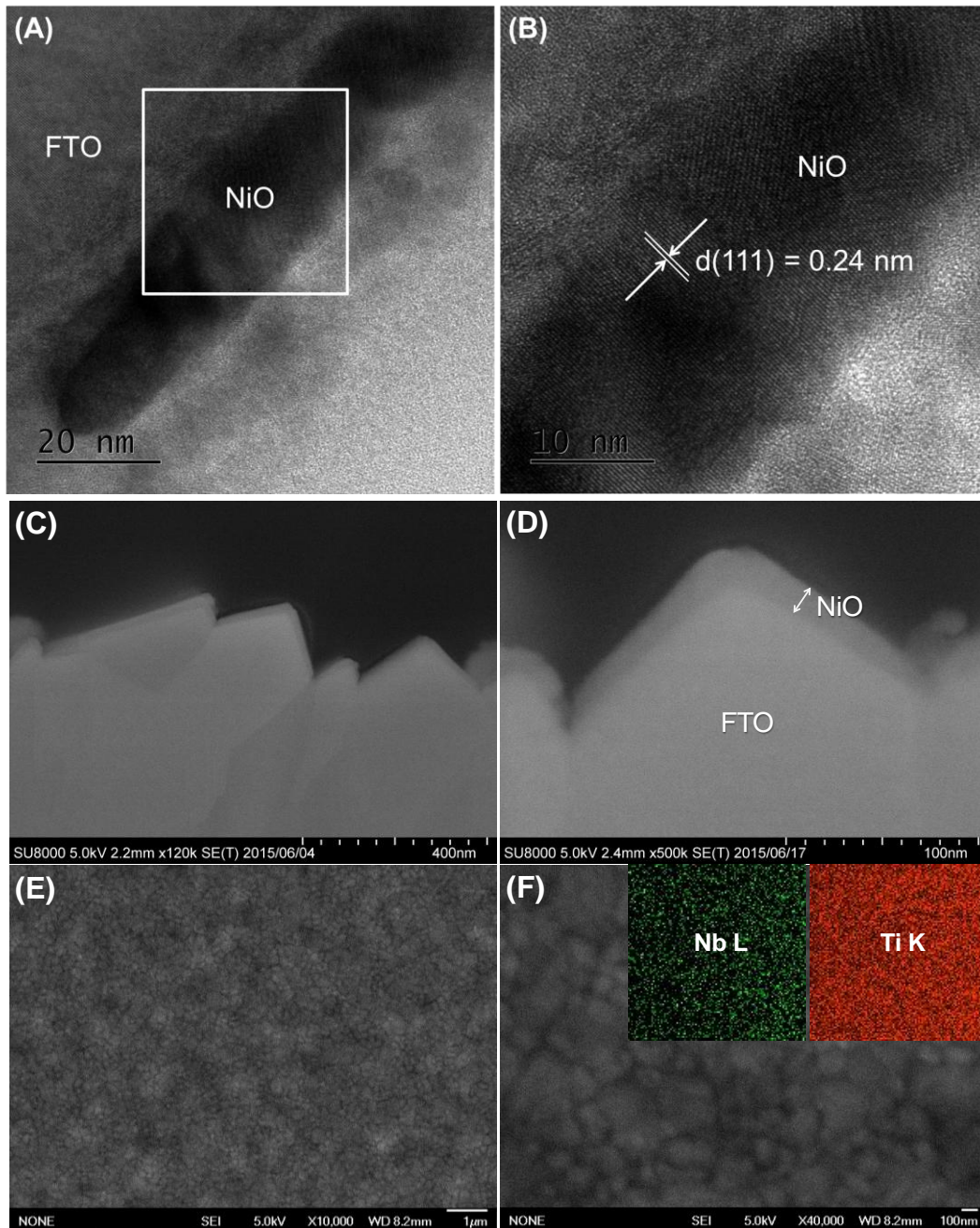
(A–B) Top view and (C–D) cross-sectional view of SEM images of MAPbI<sub>3</sub> film on FTO glass. The average crystal size is about 300 nm and average film thickness is about 300 nm. Therefore, the number of grain boundaries within the cross-sectional charge transport paths is small. (E–F) cross-section view SEM images of MAPbI<sub>3</sub> films coated with PCBM layer. The average thickness of PCBM is 80 nm. The surface roughness becomes lower after PCBM coating.



**Fig. S3**

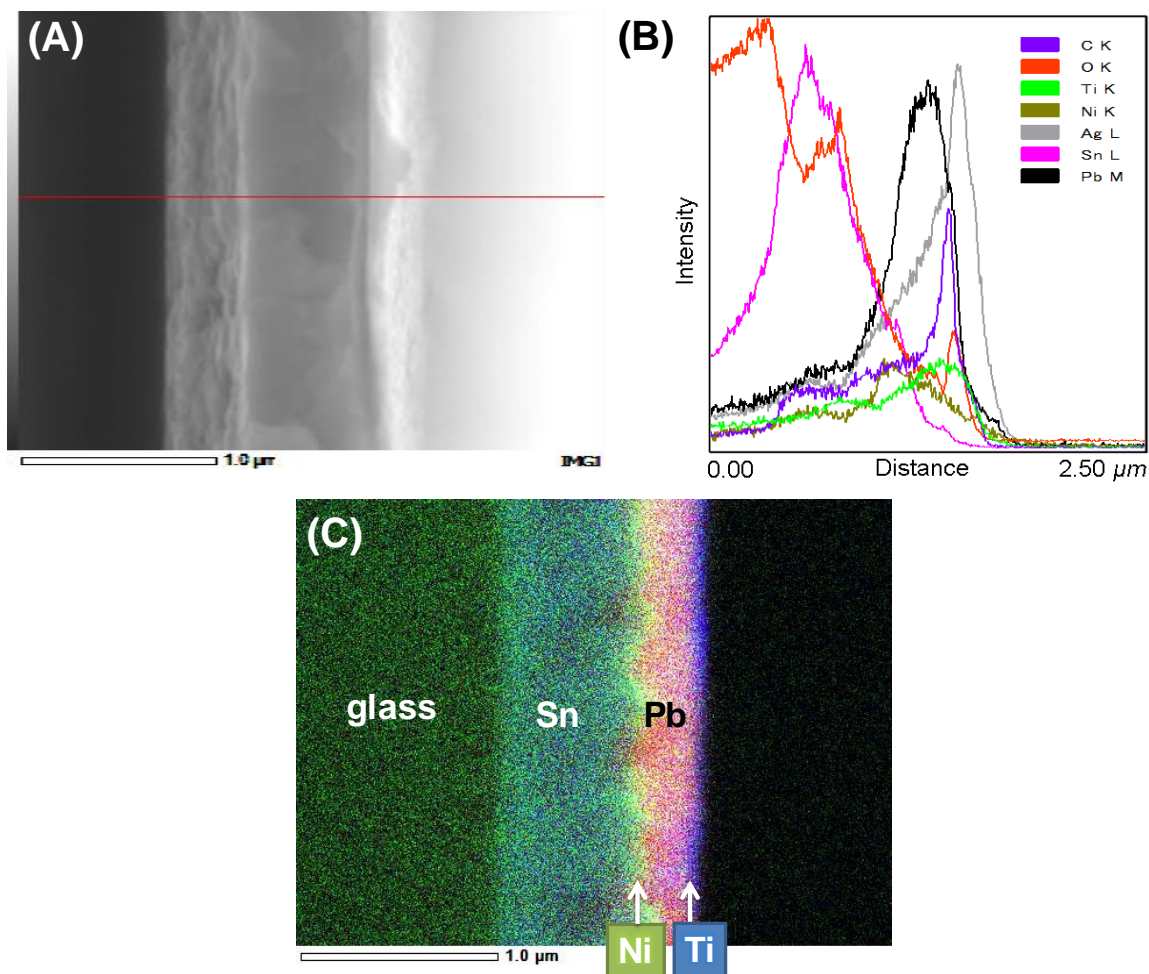
Wide survey XPS spectra of (A) NiO film, (B) NiMgLiO film, with atomic ratios of Ni: Mg: Li = 81.2: 14.5: 4.3 calculated by the relative peak area/ sensitivity factor from the narrow scan data in (C); Wide survey XPS spectra of (D) TiO<sub>x</sub> film, (E) Ti(Nb)O<sub>x</sub>, with atomic ratio of Ti: Nb = 94.7: 5.3 calculated by the relative peak area/ sensitivity factor from the narrow scan data in (F).





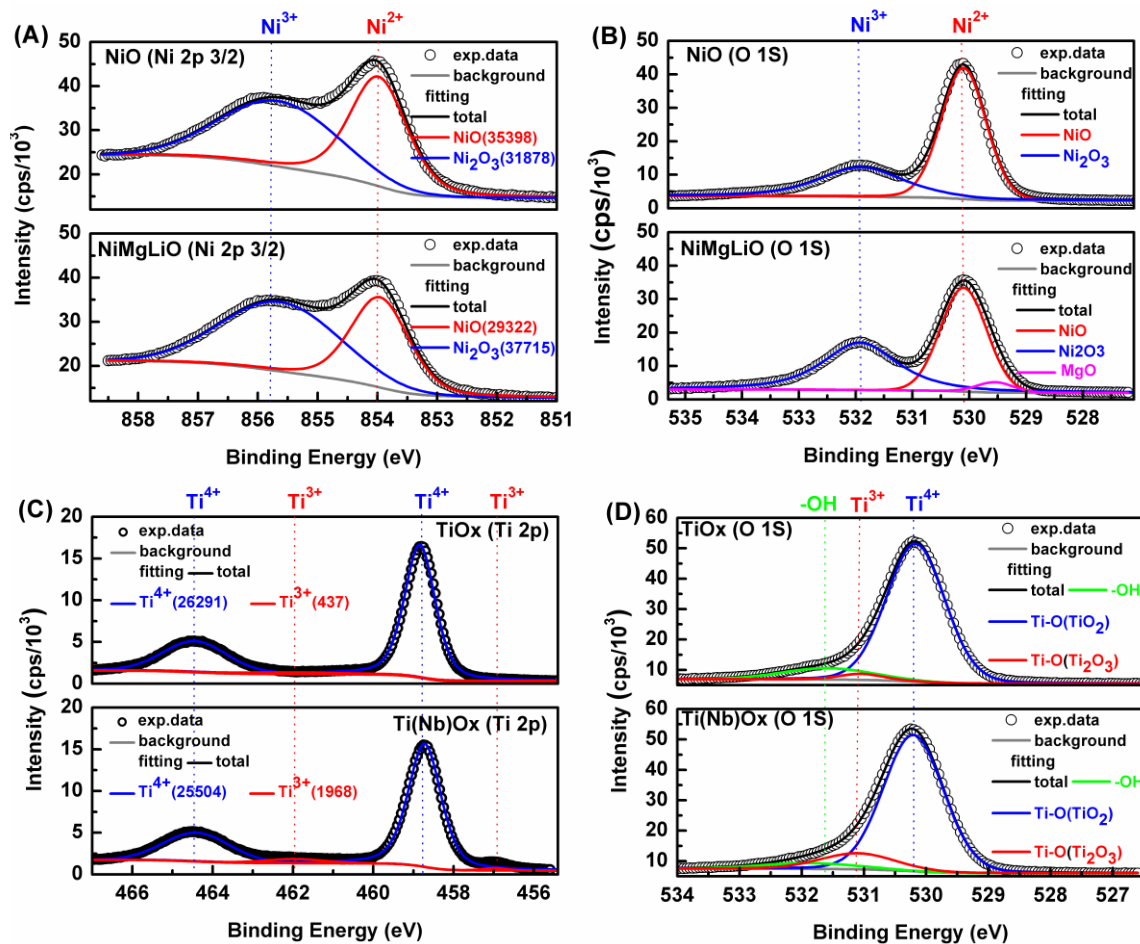
**Fig. S4**

Cross-section view of (A–B) TEM images and (C–D) backscattered electron (SE(T)) SEM images of 20 nm thick NiO films on FTO glasses; the cross sectional samples were prepared by the special FIB cutting technique; from high resolution TEM image in (B), the crystalline lattice owing to NiO could be identified; the apparent contrast between FTO and NiO coatings could be clearly identified from the SEM images, demonstrating a highly compact and uniform overage of NiO on FTO. (E–F) top view SEM images of Ti(Nb)O<sub>x</sub> coated MAPbI<sub>3</sub>/PCBM films on FTO glasses. The insets in (F) are EDX mapping results at the same SEM observed region, reflecting the uniform coating morphology and elemental distribution of the Ti(Nb)O<sub>x</sub> layer.



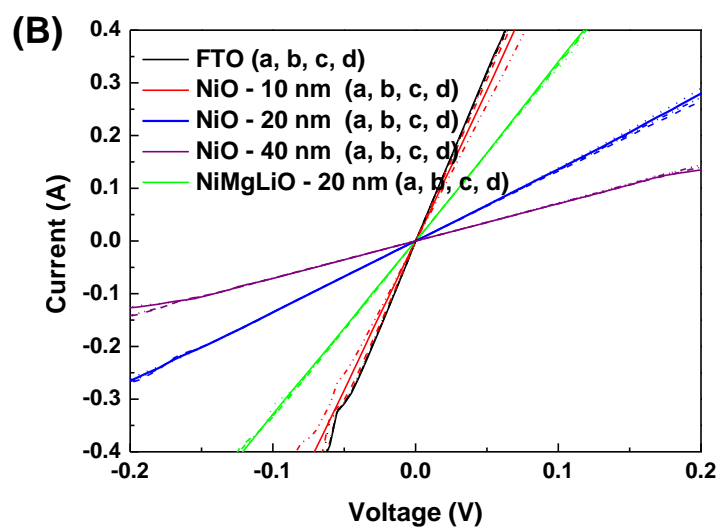
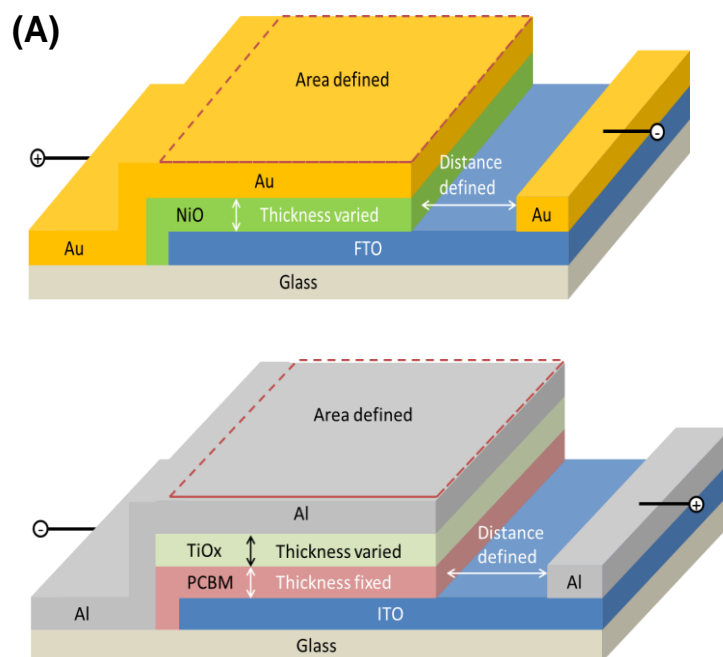
**Fig. S5**

SEM-EDX analysis results of a complete solar cell, demonstrating the layer by layer structure and NiMgLiO/Ti(Nb)O<sub>x</sub> based charge carrier extraction layers that are present in the correct locations. (A) SEM image for EDX analysis; (B) EDX linear-scan analysis result: the peak of Pb is sandwiched between Ni and Ti, and the peak of Ti is sandwiched between C and Ag; (C) the overlaid EDX mapping results of Sn (from FTO), Ni (from NiMgLiO), Pb (from perovskite) and Ti (from Ti(Nb)O<sub>x</sub>) elements.



**Fig. S6**

Deconvolution of (A) Ni 2p<sub>3/2</sub> and (B) O 1s XPS spectra of NiO/NiMgLiO films, and (C) Ti 2p and (D) O 1s XPS spectra of TiO<sub>x</sub>/Ti(Nb)O<sub>x</sub> films. The integrated areas of the fitted peaks are denoted in the corresponding brackets showing that the relative content of Ni<sup>3+</sup>/Ni<sup>2+</sup> and Ti<sup>3+</sup>/Ti<sup>4+</sup> are increased upon doping.



**(C)**

	P-type hole extraction layers' resistivity ( $10^{-1}$ ohm)			
	FTO	NiO - 10 nm	NiO - 20 nm	NiO - 40 nm
Sample-a	1.80	1.83	7.35	13.91
Sample-b	1.82	1.84	7.39	14.03
Sample-c	1.76	1.79	7.23	13.98
Sample-d	1.81	1.98	7.44	14.07

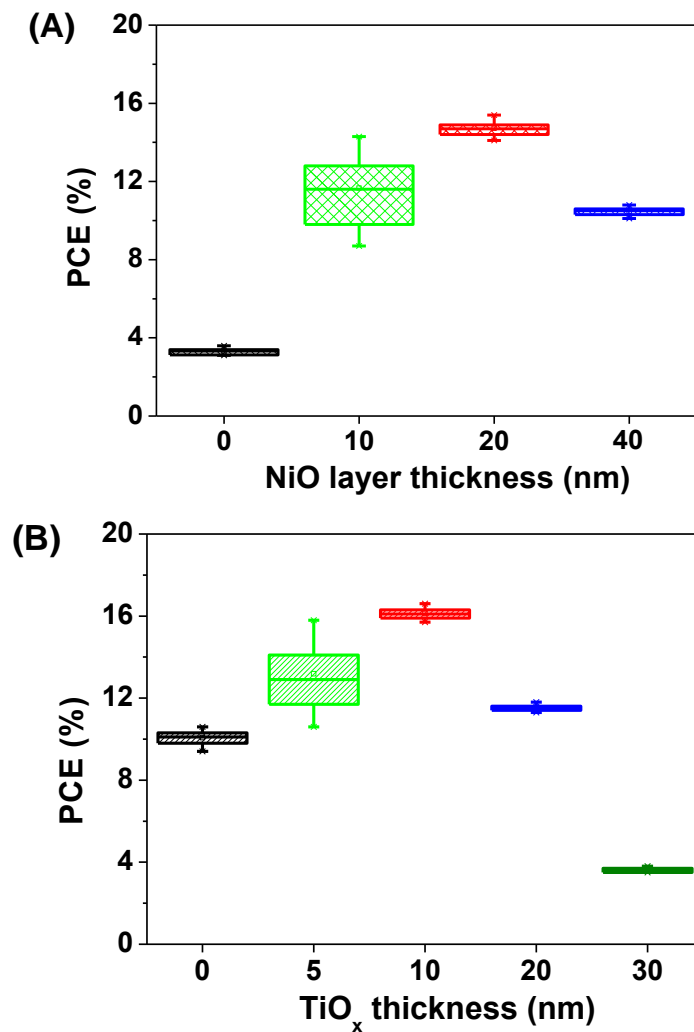
**(D)**

	N-type electron extraction layers' resistivity (ohm)			
	PCBM	TiOx - 5 nm	TiOx - 10 nm	TiOx - 20 nm
Sample-a	0.23	0.24	1.11	2.03
Sample-b	0.23	0.29	1.16	2.05
Sample-c	0.23	0.25	1.15	2.10
Sample-d	0.24	0.31	1.19	2.07

**Fig. S7**

Electrical measurement to check the pin-hole free morphology of the charge extraction layers. **(A)** The design patterns for the measurement: for NiO films, a high work function Au contact was used; while for TiO<sub>x</sub> films, the low work function Al contact was used, a PCBM under-layer was used to simulate the situation in real solar cells. **(B)** The *I*-*V* curves of different p-type film samples obtained by linear sweep voltammetry. **(C)** The calculated resistances of NiO films with different thicknesses from the slopes of the *I*-*V* curves for the parallel 4 samples on one substrate. **(D)** The calculated resistances of TiO<sub>x</sub> films with different thicknesses from the slopes of the *I*-*V* curves measured based on the design in **(A)**.

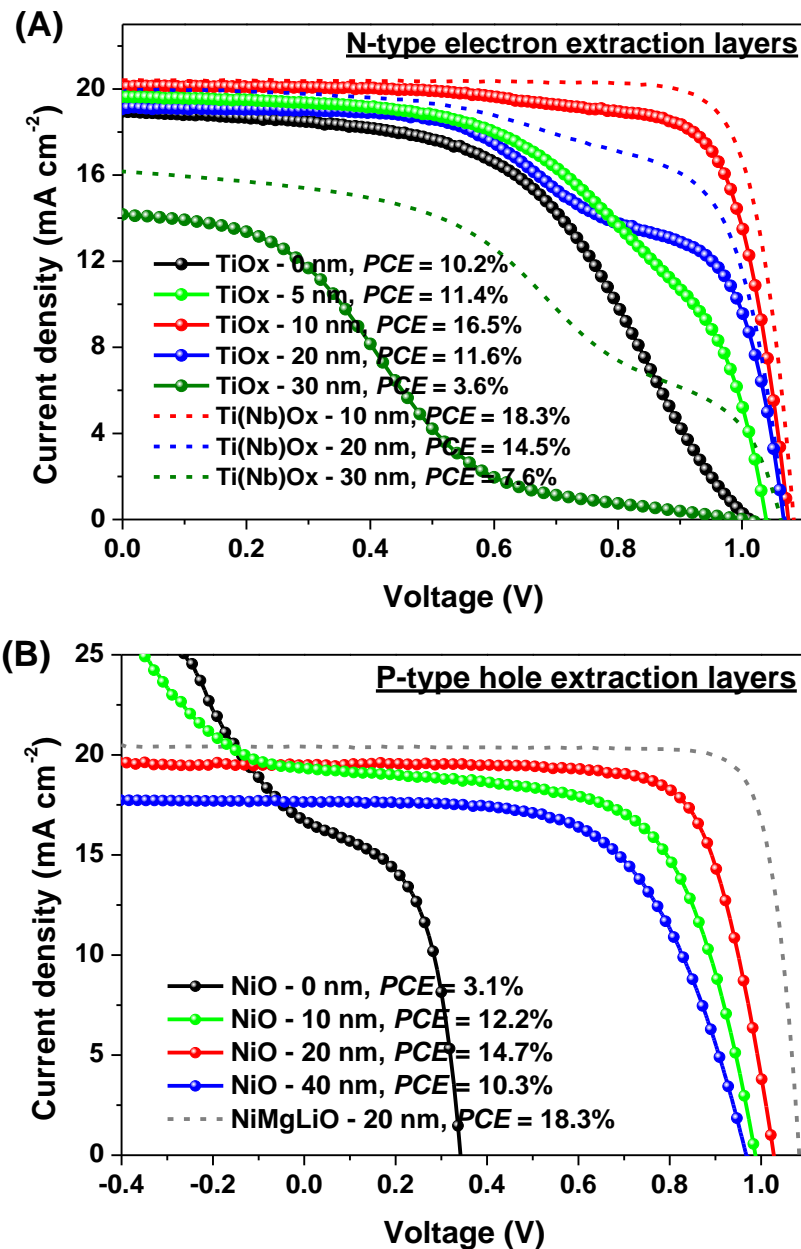
The increased resistance in comparison to bare FTO or bare PCBM are associated with charge extraction layers with different thickness. For the NiO-10 nm and TiO<sub>x</sub>-5nm films, their increased resistances are not proportional to their thickness, which is due to the presence of many pinholes inside the films leading to many short-cuts for charge leakage. This numerically electrical method is effective to judge the pin-hole density inside the blocking layers; the measurement results are consistent with the trend of the corresponding solar cells' *J*-*V* curves.



**Fig. S8**

Statistics on performance variations of the solar cells (aperture area:  $0.09 \text{ cm}^2$ ) based on the charge extraction layers with controlled thicknesses (20 cells for each condition). **(A)** The undoped NiO layer thickness = 0, 10, 20, 40 nm, with fixed  $\text{Ti(Nb)O}_x$  layer thickness of 10 nm; **(B)** The undoped  $\text{TiO}_x$  layer thickness = 0, 5, 10, 20, 30 nm with fixed NiMgLiO layer thickness of 20 nm.



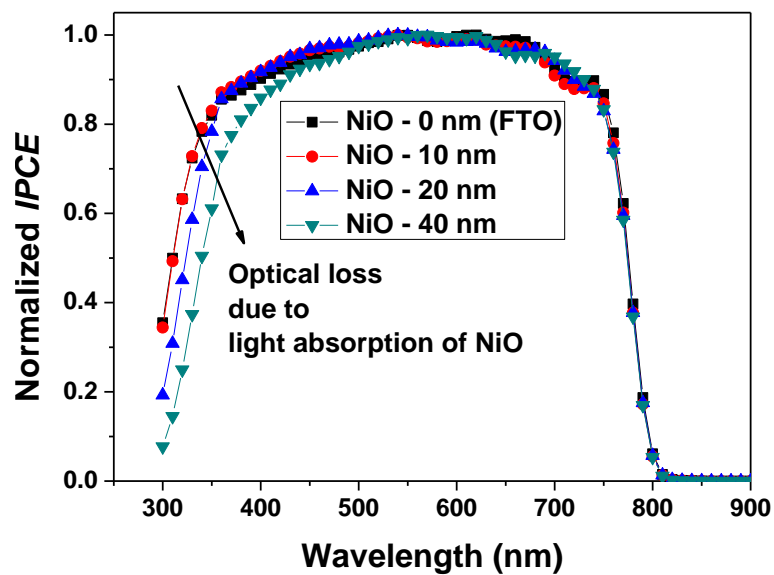


**Fig. S9**

Typical  $J$ - $V$  curves of the solar cells ( $0.09 \text{ cm}^2$ ) based on the charge carrier extraction layers with controlled thickness. **(A)** The thicknesses of TiO<sub>x</sub> films varied from 0, 5, 10, 20 to 30 nm and Ti(Nb)O<sub>x</sub> films varied from 10, 20 to 30 nm, with fixed 20 nm NiMgLiO. The largely decreased current at the high voltage range of 0.7–1.0 V is due to poor electron extraction by the TiO<sub>x</sub> layer: for 5 nm TiO<sub>x</sub> film, the decreased current is due to insufficient coverage; while for TiO<sub>x</sub> films thicker than 10 nm, largely increased bulk resistance of TiO<sub>x</sub> becomes a hindering factor for electron transport; with the same thickness of 10, 20 and 30 nm, the doped Ti(Nb)O<sub>x</sub> films performed better than the undoped TiO<sub>x</sub> films due to improved electron extraction, meanwhile the thickness increase of Ti(Nb)O<sub>x</sub> films resulted in smaller performance decline. **(B)** The thicknesses

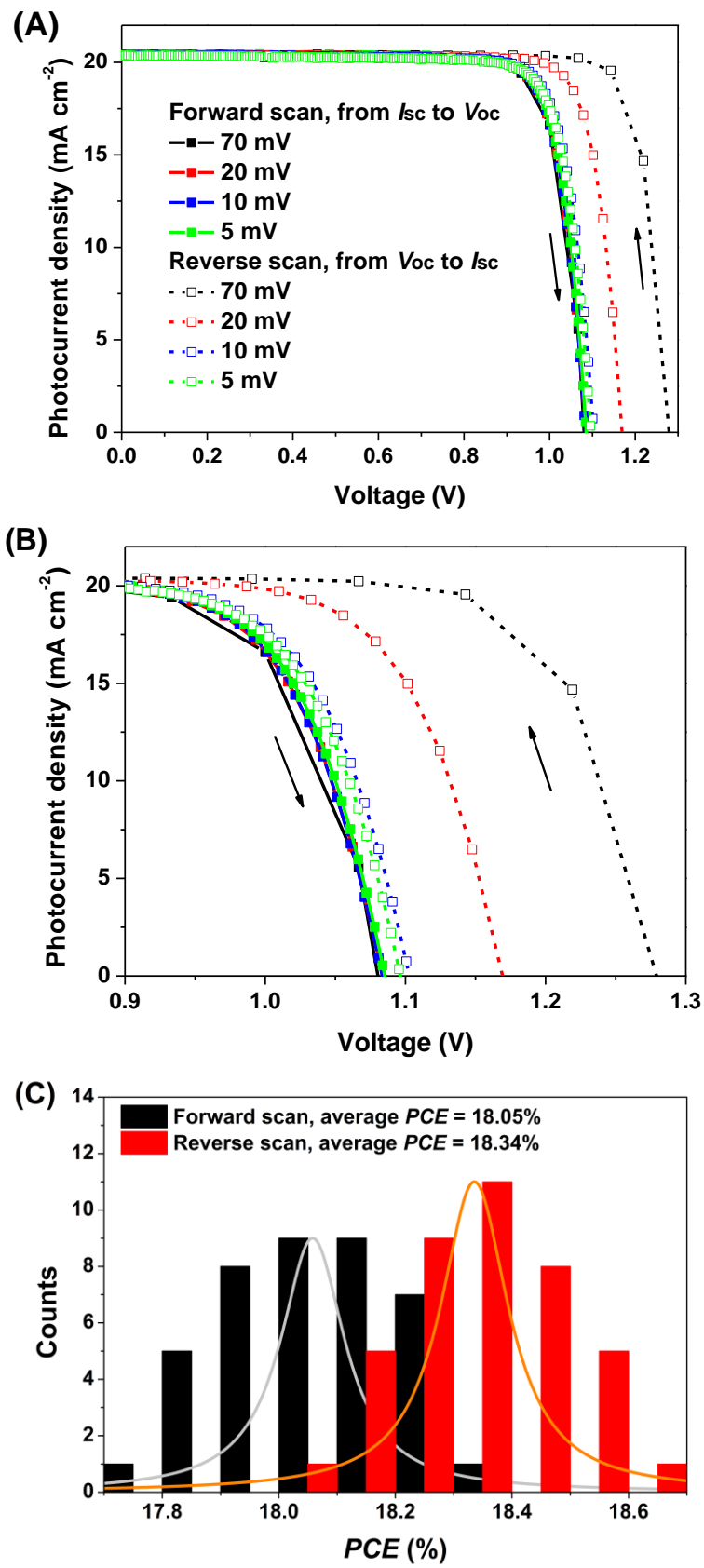
of NiO films varied from 0, 10, 20 to 40 nm and NiMgLiO film of 20 nm, with fixed 10 nm Ti(Nb)O<sub>x</sub>. The largely increased currents at the bias region NiO–0 nm and NiO–10 nm samples are due to the short-cuts between perovskite and FTO, which is consistent with the TiO<sub>2</sub> based PSC with porous blocking layer reported in (50).

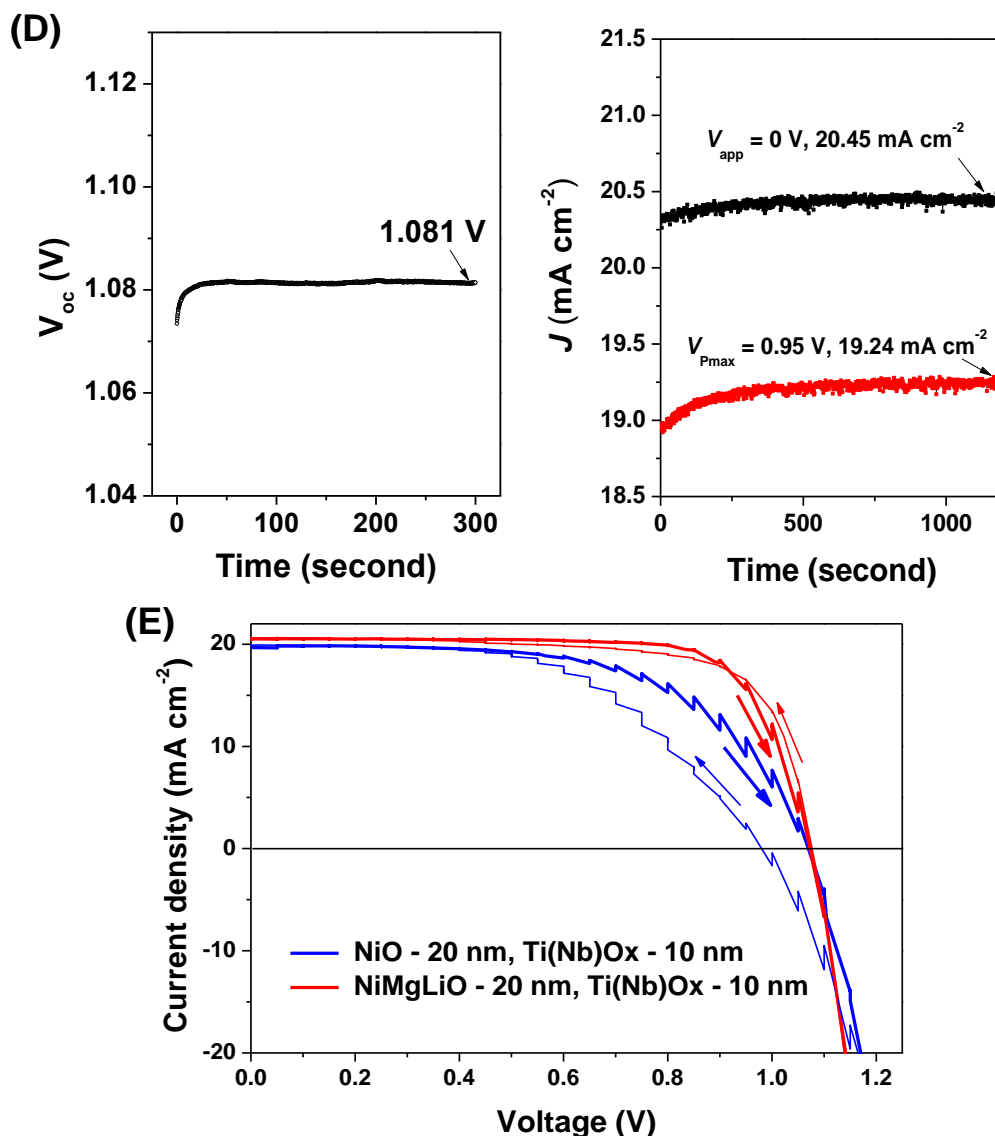




**Fig. S10**

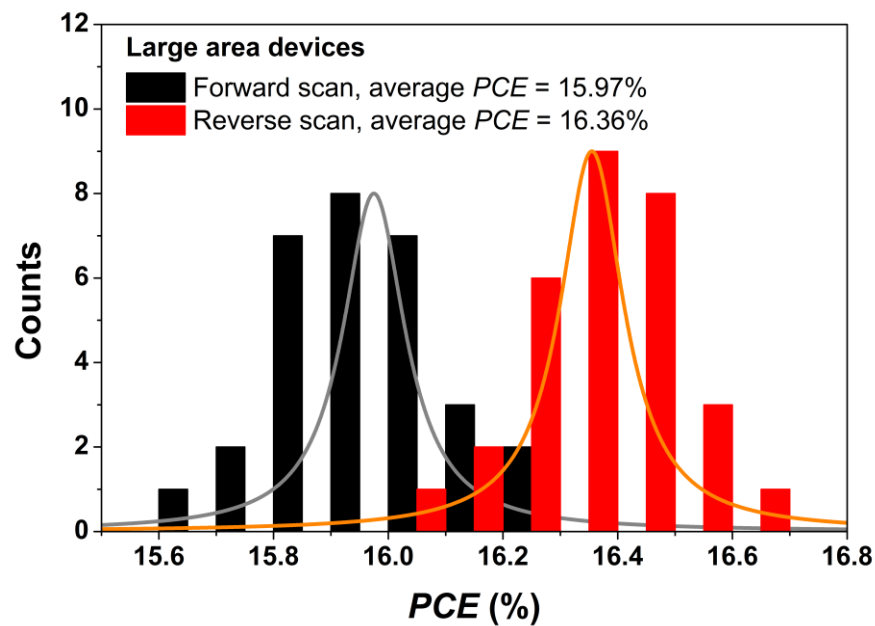
Normalized *IPCE* spectra of the solar cells based on NiO charge carrier extraction layers with different thicknesses. Optical loss in the short wavelength region (<500 nm) increases as the NiO film thickness increasing, leading to lower *IPCE* in that region.





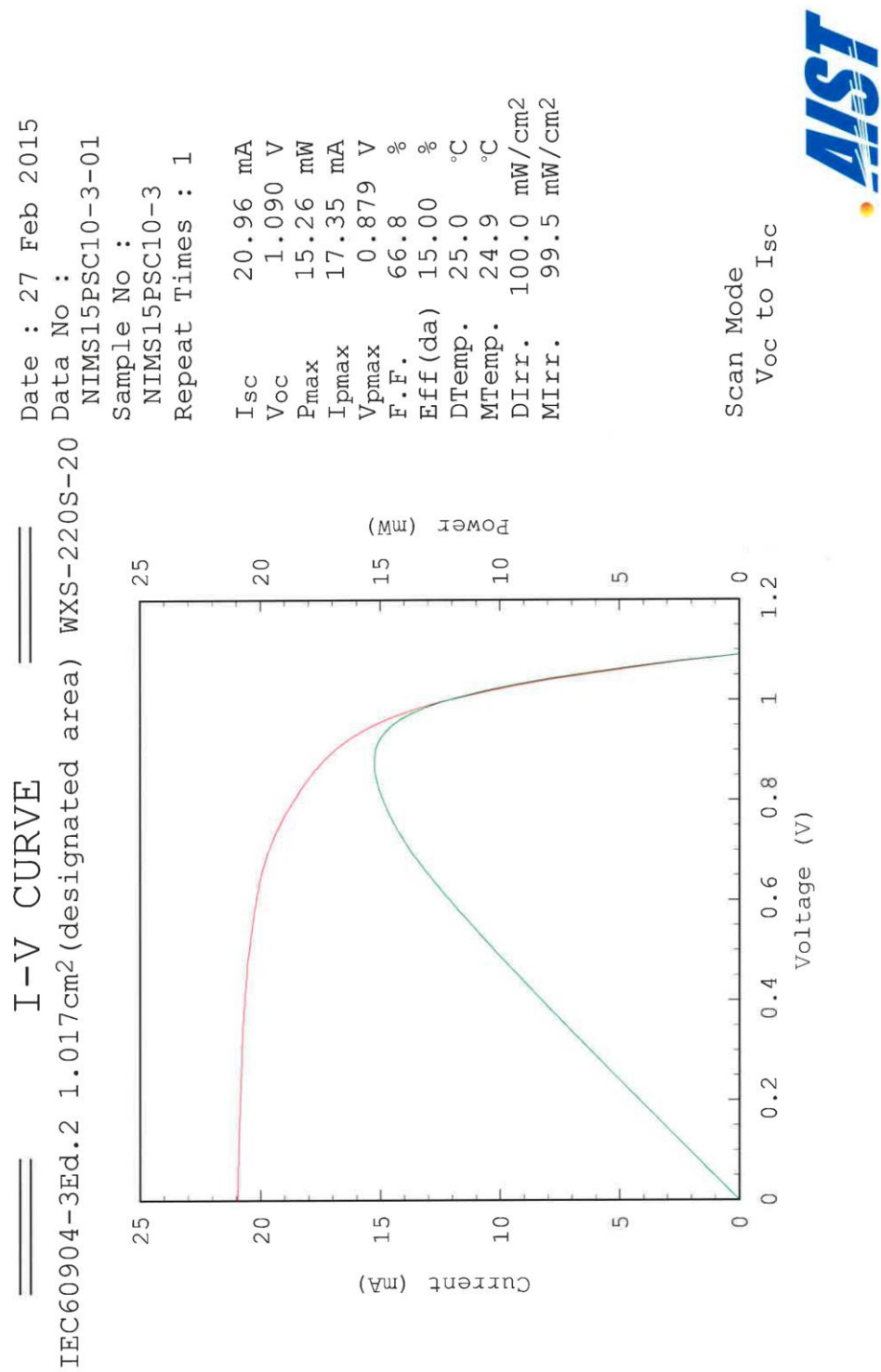
**Fig. S11**

(A-B) The  $J$ - $V$  curves of the PSC ( $0.09 \text{ cm}^2$ ) with optimal charge extraction layers, obtained at different scan directions and step widths (5 mV–70 mV), with fixed delay time of 50 ms. (B) is the enlargement of (A) in bias region of 0.9–1.3 V. (C) A histogram comparing the difference in the  $PCE$ s obtained from scanning in the forward and reverse bias directions at the step width of 10 mV, based on a set of data measured with 40 different devices made from several batches. (D) The steady state  $V_{oc}$ ,  $J_{sc}$  and  $P_{max}$  outputs of the solar cell in (A). The stable  $P_{max}$  obtained at 0.95 V bias equals to 18.28%  $PCE$  of the solar cell, which is consistent with the  $J$ - $V$  curves at small step widths (5–10 mV). (E) The stepwise  $J$ - $V$  curves of the controlled solar cells with different interfacial conditions tested on an electrochemical workstation, with two opposite scan directions, step width = 50 mV and delay time = 5000 ms. The  $J$ - $V$  hysteresis phenomenon will be enlarged at this scan condition. This result illustrates the effect of the charge carrier extraction layer composition on the hysteresis behavior of the PSC.



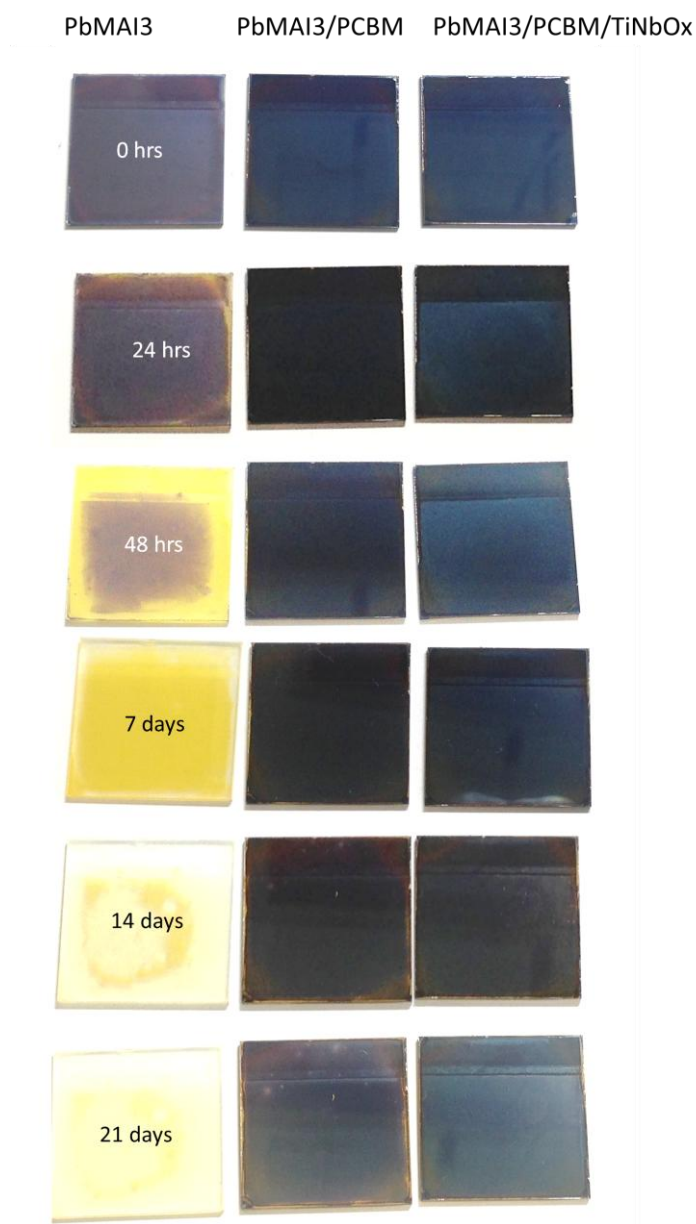
**Fig. S12**

A histogram comparing the difference in the  $PCE$ s of large area ( $1.02 \text{ cm}^2$ ) solar cells obtained from scanning in the forward and reverse bias directions, at the step width of 10 mV and delay time of 50 ms, based on a set of data measured with 30 different devices made from several batches. The average  $PCE$ s obtained from the two opposite scan directions are 15.97% and 16.36%, respectively.



**Fig. S13**

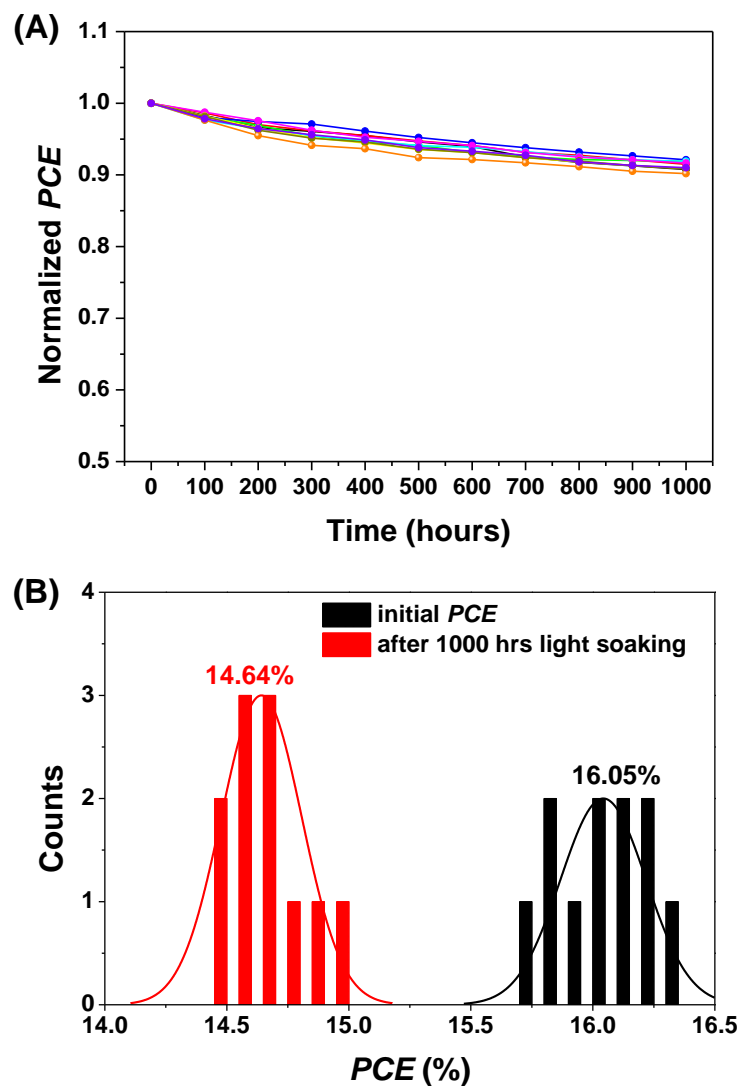
The certified results from AIST with calibrated cell size of 1.017 cm<sup>2</sup>.



**Fig. S14**

Comparison of the perovskite layer degradation speeds (by monitoring color change from deep brown to light yellow and even transparent) without and with PCBM, Ti(Nb)O<sub>x</sub> protection layers. The films were kept in the ambient air and room light condition.

Bare perovskite film decomposed evidently after 24 hours, while PCBM and PCBM/Ti(Nb)O<sub>x</sub> covered perovskite films require at least two weeks to show first signs of decomposition. After 2–3 weeks, the different decomposition speeds between PCBM and PCBM/Ti(Nb)O<sub>x</sub> covered perovskite films are apparent from their color difference.



**Fig. S15**

A batch of 10 large cells ( $1.02 \text{ cm}^2$ ) with good sealing show similar *PCEs* degradation by <10%, after exposing to full sunlight from a solar simulator for 1000 hrs at short-circuit condition: (A) the normalized *PCEs*, (B) a histogram comparing the difference in the *PCEs* before and after 1000 hrs light soaking.

**Table S1.**

Summary on the landmark progresses on efficiency evolution of perovskite solar cells, the reported large size solar cells are included for better comparison to the present work.

Cell Structure	PCE (%)	Area (cm <sup>2</sup> )	Hysteresis (Yes/No)	Certification (Yes/No)	Reference
c-TiO <sub>2</sub> /meso-Al <sub>2</sub> O <sub>3</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	10.9	0.09	Yes	No	<i>Science</i> , <b>2012</b> , 338, 643
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	14.14	0.209	Not stated	Yes, from Newport	<i>Nature</i> , <b>2013</b> , 499, 316
c-TiO <sub>2</sub> /MAPbI <sub>3-x</sub> Cl <sub>x</sub> /Spiro-OMeTAD	15.4	0.09	Yes	No	<i>Nature</i> , <b>2013</b> , 501, 395
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPb(I <sub>1-x</sub> Br <sub>x</sub> ) <sub>3</sub> /PTAA	16.2	0.16	No	Yes, from Newport	<i>Nat. Mater.</i> , <b>2014</b> , 13, 897
c-Y-TiO <sub>2</sub> /MAPbI <sub>3-x</sub> Cl <sub>x</sub> /Spiro-OMeTAD	19.3	0.1	Yes	No	<i>Science</i> , <b>2014</b> , 345, 524
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /meso-ZrO <sub>2</sub> /(5-AVA) <sub>x</sub> (MA) <sub>1-x</sub> PbI <sub>3</sub> /Carbon	12.8	0.07	No	Yes, from Newport	<i>Science</i> , <b>2014</b> , 345, 295
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /(FAPbI <sub>3</sub> ) <sub>1-x</sub> (MAPbBr <sub>3</sub> ) <sub>x</sub> /PTAA	17.9	0.09	No	Yes, from Newport	<i>Nature</i> , <b>2015</b> , 517, 476
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /(FAPbI <sub>3</sub> ) <sub>1-x</sub> (MAPbBr <sub>3</sub> ) <sub>x</sub> /PTAA	20.1	0.09	No	Yes, from Newport	<i>Science</i> , <b>2015</b> , 348, 1234
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	19.7	0.125	Not stated	No	<i>JACS</i> , <b>2015</b> , 137, 869
PEDOT: PSS/MAPbI <sub>3-x</sub> Cl <sub>x</sub> /PCBM	18.0	0.035	No	No	<i>Science</i> , <b>2015</b> , 347, 522
c-NiO <sub>x</sub> /MAPbI <sub>3</sub> /PCBM/LiF	17.3	0.04	Not stated	No	<i>Adv. Mater.</i> , <b>2015</b> , 27, 4013
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	12.6	1	Not stated	No	<i>Nano Energy</i> , <b>2015</b> , 15, 670
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	13.3 10.4 4.3	0.1 10.1 100	Not stated	No	<i>J. Power. S.</i> , <b>2015</b> , 277, 286
c-TiO <sub>2</sub> /meso-TiO <sub>2</sub> /MAPbI <sub>3</sub> /Spiro-OMeTAD	10.5	10.8	Not stated	No	<i>ACS Nano</i> , <b>2015</b> , 9, 8420
PEDOT: PSS/MAPbI <sub>3</sub> /PCBM/C <sub>60</sub>	10.9	0.95	No	No	<i>Adv. Energy Mater.</i> , <b>2014</b> , 4, 1400345
NiMgLiO/MAPbI <sub>3</sub> /PCBM/Ti(Nb)O <sub>x</sub>	15.0	1.017	No	Yes, from AIST	<i>The present work</i>



**Table S2.**

Photovoltaic performance metrics of the typical devices employing various combinations of charge carrier extraction layers with their  $J$ – $V$  curves shown in Fig. 3A. The  $R_{sh}$  and  $R_s$  are obtained by linear fitting of the slopes for the  $J$ – $V$  curves near the short-circuit condition and open-circuit condition, respectively.

<b>Interfacial Conditions</b>	<b><math>J_{sc}(\text{mA cm}^{-2})</math></b>	<b><math>V_{oc}(\text{V})</math></b>	<b><math>FF</math></b>	<b><math>PCE(\%)</math></b>	<b><math>R_{sh}(\text{Ohm cm}^2)</math></b>	<b><math>R_s(\text{Ohm cm}^2)</math></b>
NiO/TiO <sub>x</sub>	18.7	1.036	0.640	12.4	1978	7.38
NiO/Ti(Nb)O <sub>x</sub>	19.2	1.047	0.731	14.7	3597	5.43
NiMgLiO/TiO <sub>x</sub>	20.0	1.074	0.768	16.5	2237	3.59
NiMgLiO/Ti(Nb)O <sub>x</sub>	20.4	1.083	0.827	18.3	5605	2.45

**Table S3.**

Efficiency parameters of the perovskite solar cell from  $J$ - $V$  curves obtained at difference scan conditions. The corresponding  $J$ - $V$  curves are shown in Fig. S11A.

Scan Direction	Step Width (mV)	$J_{sc}$ (mA cm <sup>-2</sup> )	$V_{oc}$ (V)	$FF$	$PCE$ (%)
From $V_{oc}$ to $I_{sc}$	70	20.429	1.273	0.859	22.35
	20	20.427	1.167	0.836	19.92
	10	20.401	1.103	0.825	18.56
	5	20.387	1.096	0.823	18.40
From $I_{sc}$ to $V_{oc}$	70	20.399	1.080	0.823	18.14
	20	20.431	1.084	0.823	18.23
	10	20.411	1.083	0.827	18.29
	5	20.418	1.085	0.830	18.39

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