

DOI: 10.1126/science.1230767 Science **339**, 554 (2013); Shau-Yu Lan et al. **A Clock Directly Linking Time to a Particle's Mass**

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<http://www.sciencemag.org/cgi/collection/physics> **Physics** This article appears in the following **subject collections:** interrelationships between diet, microbiota, and many facets of host physiology can be explored in detail in these "personalized" gnotobiotic mouse models. These models may be useful for developing new and more effective approaches for treatment and/or prevention. In addition, studies of other forms of malnutrition that take an approach analogous to that described here could also provide insights about the contribution of the gut microbiome to this global health problem.

References and Notes

- 1. United Nations (UN) Inter-Agency Group for Child Mortality Estimation, Levels & Trends in Child Mortality Report (2011); www.childinfo.org/files/Child_Mortality_ Report_2011.pdf.
- 2. WHO Multicentre Growth Reference Study Group, Acta Paediatr. Suppl. 450, 76 (2006).
- 3. WHO, UN Children's Fund, WHO Child Growth Standards and the Identification of Severe Acute Malnutrition in Infants and Children (2009); www.who.int/nutrition/publications/ severemalnutrition/9789241598163/en/index.html.
- 4. C. D. Williams, B. M. Oxon, H. Lond, Lancet 226, 1151 (1935).
- 5. T. Ahmed, S. Rahman, A. Cravioto, Indian J. Med. Res. 130, 651 (2009).
- 6. C. Gopalan, in Calorie Deficiencies and Protein Deficiencies: Kwashiorkor and Marasmus: Evolution and Distinguishing Features, R. A. McCance, E. M. Widdowson, Eds. (Churchill, London, 1968), pp. 48–58.
- 7. M. H. Golden, Lancet 319, 1261 (1982). 8. H. Ciliberto et al., BMJ 330, 1109 (2005).
- 9. C. A. Lin et al., J. Pediatr. Gastroenterol. Nutr. 44, 487 (2007).
- 10. T. Yatsunenko et al., Nature 486, 222 (2012).
- 11. R. E. Black et al., Lancet 371, 243 (2008).
- 12. WHO, World Food Programme, UN System Standing Committee on Nutrition, UN Children's Fund, Community-Based Management of Severe Acute Malnutrition (2007); www.who.int/nutrition/topics/ statement_commbased_malnutrition/en/index.html.
- 13. Materials and methods are available as supplementary materials on Science Online.
- 14. L. Lagrone, S. Cole, A. Schondelmeyer, K. Maleta, M. J. Manary, Ann. Trop. Paediatr. 30, 103 (2010).
- 15. J. Liu et al., J. Clin. Microbiol. 50, 98 (2012).
- 16. M. Taniuchi et al., Diagn. Microbiol. Infect. Dis. 71, 386 (2011).
- 17. M. Taniuchi et al., Am. J. Trop. Med. Hyg. 84, 332 (2011).
- 18. M. Taniuchi et al., Diagn. Microbiol. Infect. Dis. 73, 121 (2012).
- 19. S. Devkota et al., Nature 487, 104 (2012).
- 20. N. Crum-Cianflone, Am. J. Med. Sci. 337, 480 (2009).
- 21. T. Itoh, Y. Fujimoto, Y. Kawai, T. Toba, T. Saito, Lett. Appl. Microbiol. 21, 137 (1995).
- 22. M. F. Fernández, S. Boris, C. Barbés, J. Appl. Microbiol. 94, 449 (2003).
- 23. Y. Kato-Mori et al., J. Med. Food 13, 1460 (2010).
- 24. A. A. Salyers, S. E. West, J. R. Vercellotti, T. D. Wilkins,
- Appl. Environ. Microbiol. 34, 529 (1977). 25. H. Sokol et al., Proc. Natl. Acad. Sci. U.S.A. 105, 16731 (2008).
- 26. O. Beckonert et al., Nat. Protoc. 2, 2692 (2007).
- 27. H. Laue, K. Denger, A. M. Cook, Appl. Environ. Microbiol. 63, 2016 (1997).
- 28. J. C. Edozien, E. J. Phillips, W. R. F. Collis, Lancet 275, 615 (1960).
- 29. R. G. Whitehead, R. F. Dean, Am. J. Clin. Nutr. 14, 313 (1964).
- 30. S. Awwaad, E. A. Eisa, M. El-Essawy, J. Trop. Med. Hyg. 65, 179 (1962).
- 31. G. Arroyave, D. Wilson, C. DeFunes, M. Béhar, Am. J. Clin. Nutr. 11, 517 (1962).
- 32. T. R. Ittyerah, S. M. Pereira, M. E. Dumm, Am. J. Clin. Nutr. 17, 11 (1965).
- 33. T. R. Ittyerah, Clin. Chim. Acta 25, 365 (1969).
- 34. D. H. Baker, Prog. Food Nutr. Sci. 10, 133 (1986).
- 35. N. Orentreich, J. R. Matias, A. DeFelice, J. A. Zimmerman, J. Nutr. 123, 269 (1993).
- 36. M. Mori, S. Manabe, K. Uenishi, S. Sakamoto, Tokushima J. Exp. Med. 40, 35 (1993).
- 37. J. K. Nicholson, J. A. Timbrell, P. J. Sadler, Mol. Pharmacol. 27, 644 (1985).

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Supplementary Materials

www.sciencemag.org/cgi/content/full/339/6119/548/DC1 Materials and Methods Supplementary Text Figs. S1 to S15 Tables S1 to S10 References (38–54)

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REPORTS

to a Particle's Mass

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Historically, time measurements have been based on oscillation frequencies in systems of particles, from the motion of celestial bodies to atomic transitions. Relativity and quantum mechanics show that even a single particle of mass m determines a Compton frequency $\omega_0 = mc^2/\hbar$, where c is the speed of link-
linkt and \hbar is Planck's constant h divided by 2π . A clock referenced to ω_2 would enable bigh-precisio light and \hbar is Planck's constant h divided by 2π . A clock referenced to ω_0 would enable high-precision mass measurements and a fundamental definition of the second. We demonstrate such a clock using an optical frequency comb to self-reference a Ramsey-Bordé atom interferometer and synchronize an oscillator at a subharmonic of ω_0 . This directly demonstrates the connection between time and mass. It allows measurement of microscopic masses with 4×10^{-9} accuracy in the proposed revision to SI units. Together with the Avogadro project, it yields calibrated kilograms.

A particle with mass-energy $E = mc^2$ is

compton frequency $\omega_0 = mc^2/\hbar$ in the

particle's rest frame where c is the speed of light represented by a wave oscillating at the particle's rest frame, where c is the speed of light

and \hbar is Planck's constant h divided by $2\pi (1)$. This is the basis of de Broglie's theory of matter waves (2) and underpins modern quantum mechanics and field theory: The time evolution of states is given by wave equations whose planewave solutions are proportional to

$$
e^{-i\phi} = \exp(-ip_{\mu}x^{\mu}/\hbar) = \exp(-i\omega_0\tau) \qquad (1)
$$

where $p_{\mu} = (-m\gamma, m\gamma \nu)$ and x^{μ} are the momen-
turn and position four-vector $\tau = t/\nu$ is the proper tum and position four-vector, $\tau = t/\gamma$ is the proper time, γ is the Lorentz factor, and ν and t are the laboratory-frame velocity and time. Much has been theorized about the physical reality of quantum states as "oscillators" $(3-7)$, but surprisingly few experiments have been proposed to address this topic (8) . Here, we directly address a consequence of Eq. 1 that has deep physical and perhaps even cosmological implications: Because the oscillations of a wave packet accumulate phase $\omega_0 \tau$ just like a clock following the same

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trajectory, it has been suggested that the passage of time is observable as soon as the universe contains massive particles (1) . It ought to be possible to build a Compton clock: a clock referenced to the mass of a single particle.

Advances in timekeeping (9–12) have followed the use of references with higher frequency, increased quality factor Q , and lowered instability from systematic influences. The Compton frequency of a stable particle is high ($\omega_0/2\pi = 3 \times$ 10^{25} Hz for a cesium atom), has virtually infinite quality factor $Q = \omega_0/\Gamma$, where $1/\Gamma$ is the particle's 1/e lifetime, and is one of the most stable quantities in nature. In principle, a Comptonfrequency clock could be built by annihilating a particle-antiparticle pair and counting the frequencies of the generated photons. However, this frequency is far beyond modern counting techniques; to access it, we require a method of dividing it into a technically accessible range.

Here, we present a Compton clock by combining an atom interferometer with an optical frequency comb. In a Ramsey-Bordé atom interferometer (Fig. 1) $(7, 13)$, a quantum wave packet initially at rest interacts with a pulse of two counterpropagating laser beams having laboratoryframe frequencies of ω_{+} and wavenumbers of k_{\pm} , respectively. The wave packet absorbs *n* photons from the first beam while being stimulated to emit n photons into the second beam, without modifying its internal state through multiphoton Bragg diffraction (14, 15). By appropriate choice of the laser intensity and pulse duration, the process occurs with a probability of 50%, splitting the wave packet. While the first partial wave packet remains at rest, the second moves away with a laboratory-frame momentum $p = n\hbar(k_{+} + k_{-})$. It is convenient to introduce a rapidity parameter ρ so that $p = mcsinh(\rho)$ and the Lorentz factor $\gamma = [1 - v^2/c^2]^{-1/2} = \cosh(\rho)$. Energy-momentum con-
servation (the Bragg condition) provides the relation servation (the Bragg condition) provides the relation

$$
\omega_{\pm} = \omega_{\mathrm{L}} e^{\pm \rho/2} \tag{2}
$$

where $\omega_L = \omega_0 \sinh(\rho/2)/n$. [In the nonrelativistic limit, $\omega_L \approx (\omega_+ + \omega_-)/2 \equiv kc$ and $\omega_{\pm} \approx \omega_L (1 \pm n\hbar k/2)$ mc).] Interaction with four laser pulses (Fig. 1) brings the wave packets together for interference.

Fig. 1. Paths of the matter waves in a Ramsey-Bordé interferometer versus time; red arrows denote laser beams. Additional paths that do not interfere are not shown. The diagram is drawn in a freely falling inertial frame, and does not show the free fall of the atoms.

The probability of detecting the particle at the outputs (Fig. 1) is given by $cos^2(Δφ/2)$, where $Δφ$
is the relative phase of the wave packets. In a is the relative phase of the wave packets. In a relativistic, semiclassical treatment, Eq. 1 shows that the wave packets acquire a relative phase of $\Delta \varphi_{\text{free}} = \omega_0 [\tau^{(1)} - \tau^{(2)}] = 2\omega_0 T(\gamma^{-1} - 1)$ (non-
relativistically $\Delta \omega_c \approx -4n^2\hbar k^2 T/m$) during their relativistically, $\Delta \varphi_{\text{free}} \approx -4n^2 \hbar k^2 T/m$ during their
free evolution between the laser pulses (3) where free evolution between the laser pulses (3) , where difference $\tau^{(1)} - \tau^{(2)}$ is related to the laboratory-
frame time by a factor $\gamma^{-1} - 1$, which is determined $\tau^{(1,2)}$ are the proper times on paths 1 and 2. Their frame time by a factor $\gamma^{-1} - 1$, which is determined by the laser frequencies by Eq. 2 and can be used to divide the Compton frequency into a technically accessible range. Synchronizing an oscillator to the divided Compton frequency is accomplished as follows: In the experiment, this oscillator sets the frequency difference $\omega_m \equiv \omega_+ - \omega_$ between the counterpropagating laser beams. Whenever the particle absorbs a photon, the phase of the photon is added to the wave packet's phase and subtracted for stimulated emission. It can be shown (16) that the sum $\Delta\varphi_{\text{laser}}$ of these phases (eq. S2) cancels the free evolution phase $\Delta\varphi$ _{free}

$$
\omega_{\rm m} = \frac{2n\omega_{\rm L}^2}{\omega_0} \tag{3}
$$

The phase cancellation $\Delta \varphi = \Delta \varphi_{\text{laser}} + \Delta \varphi_{\text{free}} = 0$ is verified by observing the interference fringes of the particle at the interferometer output and maintained by feedback to the oscillator.

exactly if, and only if, ω_{+} satisfy Eq. 2, so that

Ramsey-Bordé interferometers are routinely used to measure \hbar/m (17), but such measurements of \hbar/m are not themselves clocks because they are reliant upon an external frequency reference to measure ω_L . We use an optical frequency comb to set $\omega_L = N \omega_m$, eliminating the laser frequency as a free parameter. Combined with Eq. 3, we obtain

$$
\omega_{\rm m} = \frac{\omega_0}{(2nN^2)}\tag{4}
$$

which is an exact result in this relativistic and semiclassical picture, showing that ω_m is determined solely by the Compton frequency. Corrections from nonclassical paths are negligible here. We are free to choose any value for n and N and lock the oscillator directly to the Compton frequency or a subharmonic without changing the underlying physics. For example, $n = 2$ and $N=$ underlying physics. For example, $n = 2$ and $N = 1/2$ yields $\omega_0 = \omega_m = 2\omega_L$ and $\omega_{\pm} = \omega_L(1 + \sqrt{2})^{\pm 1}$.
In practice, N is chosen so that ω_{\pm} and ω_{\pm} are In practice, N is chosen so that ω_L and ω_m are accessible to existing technologies. The experiment is a Compton clock by virtue of the fact that it self-references the laser frequency through the optical frequency comb, which ensures that the output frequency ω_m is fully determined by the Compton frequency and known ratios. It is independent of any external standards. The selfreferenced atom interferometer serves as a clockwork that links the frequency of our master oscillator to the Compton frequency of the matter wave packets.

Our experiment uses cesium-133 atoms because they are amenable to the techniques of atom optics and are electrically neutral, which reduces systematic effects. Because their internal structure is not relevant to $\omega_{\rm m}$, the atoms approximate noninteracting point masses. Our atom interferometer has been described in detail elsewhere (15) . The clock's operation is illustrated in Fig. 2. A voltage-controlled 10-MHz crystal oscillator O1 [FTS 1050A (Symmetricom, Beverly, MA)] with frequency ω_{cryst} is the master oscillator for all

Fig. 2. Schematic. Oscillator O1 is the frequency reference for all signal generators and the optical frequency comb. The laser used to address the atom interferometer is phase-locked to the comb. Shown in the diagram is the freely falling atomic trajectories versus time for the simultaneous conjugate interferometers (the lower of which being shown in Fig. 1) used to cancel the gravity-induced phase. The phase measurement from the atom interferometer provides an error signal to O1 in order to close the feedback loop.

frequency sources, pulse generators, and lasers involved in the experiment. The frequency comb stabilizes the laser frequency to $N_{\rm c}\omega_{\rm crust}$, where $N_{\rm c}$ = 35,173,594.165 (16). A direct digital synthesizer (DDS) generates $\omega_{\rm m}/2 = N_{\rm DDS}\omega_{\rm cryst}$ at ~82 kHz, and acousto-optical modulators (AOMs) generate ω_{\pm} by applying the relation $\omega_{\pm} = \omega_{\text{L}} \pm \omega_{\text{m}}/2$ (valid in the nonrelativistic limit), where $N_{\text{DDS}} =$ 2,326,621,801,616/2⁴⁸. The factor N is thus $N =$ $N_c/N_{\text{DDS}} = 4,255,305,521.31286$. Another AOM shapes the time domain profile of the laser pulses. Our signal to noise is optimal for 5th-order Bragg diffraction $(n = 5)$ with a pulse separation time (Fig. 1) $T = 160$ ms. The atomic population at the interferometer output is determined by means of fluorescence detection and then processed to yield the interferometer phase $\Delta \varphi = \Delta \varphi_{\text{laser}} + \Delta \varphi_{\text{free}}$.

A feedback loop adjusts ω_{cryst} so as to maintain $\Delta \varphi = 0$. Its operation is demonstrated in Fig. 3: The first 10 data points are taken during normal operation. At the 11th data point, we briefly disable the feedback, so that ω_{cryst} is free-running, and the experiment performs a Ramsey-Bordé photon recoil measurement. The measured signal is then proportional to $\omega_L^2 = (N_c \omega_{\text{crys}})^2$ (Eq. 3). Increasing N_c by 100 parts per billion (ppb) increases ω_L and produces a 200-ppb increase in $\Delta\varphi$ _{free} (Fig. 3, shaded area). At the 13th data point, the feedback is restored. Because the feedback stabilizes ω_{crvst} to a value inversely proportional to N_c^2 (Eq. 4), the signal frequency decreases by 400 ppb, equilibrating 200 ppb lower than the first 10 data points (Fig. 3), which is opposite to the observed shift for a recoil measurement.

Quantitative agreement is demonstrated by comparing the rest-mass stabilized frequency ω_{crvst} with a rubidium frequency standard [SRS FS725 (Stanford Research Systems, Sunnyvale, CA)] over 6 hours (Fig. 4A). It averages to 9,999,998.127 Hz with a SD of the mean of 0.015 Hz. A χ^2 test (Fig. 4B) yields a normalized χ^2 = 1.4. Our statistical uncertainty is 0.015 Hz \times $(\chi^2)^{1/2} = 0.018$ Hz, or 1.8 ppb. Using Eq. 4 and
correcting for systematic effects (16) we obtain correcting for systematic effects (16) , we obtain the Compton frequency $\omega_0/2\pi = (2,993,486,252 \pm 1)$ 12×10^{16} Hz. The deviation from the expected values (16, 18) is -5.2 ± 4.0 ppb and is consistent with zero within 2σ . The Allan variance (Fig. 4C) is below $10^{-8}/[\tau/(1000 \text{ s})]^{1/2}$, where τ is the integration time is the integration time.

Fig. 3. Measured frequency ω_m (lower trace; left axis) as well as $\omega_m(2nN^2)$ (upper trace; right axis) shows the action of the feedback loop. During the shaded interval the feedback is disabled shows the action of the feedback loop. During the shaded interval, the feedback is disabled.

Although this accuracy is modest, similar to the first cesium atomic clocks (9), the resolution and accuracy of atom interferometry is advancing rapidly (19). Up to 100-fold improvement in resolution could be obtained by approaching the shot noise limit of our interferometer. Use of higher laser frequencies (which is possible without changing the reference particle), lighter particles, and/or longer interrogation times can possibly lead to a Compton clock that can serve as a primary time standard. The intrinsic stability and Q factor of a stable particle's Compton frequency is unmatched by any other frequency reference. Mass, frequency, time, and length could all be derived from one fundamental unit, defined by a specific particle.

The clock can also be used for the opposite purpose, measuring mass by measuring the Compton frequency. In 2011, the General Conference on Weights and Measures (CGPM-2011) considered a revision to the SI units that would assign an exact value to the Planck constant h (20). The kilogram would then be referenced to the second through the defined values of the Planck constant and the speed of light. Microscopic masses could be related to the fine structure constant $(21, 22)$ or h/M (17) ; macroscopic masses could be measured by using the Watt balance (23, 24). These methods, however, require auxiliary measurements and/or intricate theory (16). The Compton clock would provide an absolute measurement of the cesium atom's mass to an accuracy of 4.0 ppb, which is competitive with other methods and greater than 10 times more accurate than in the present SI (18) . Other microscopic masses can be related to the cesium mass by means of mass spectroscopy.

The link to macroscopic masses can be made by Avogadro spheres: silicon crystals of accurately measured volume V and lattice constant a (25). In theory, the number of atoms contained in such a crystal is $N_{at} = 8V/a^3$, where 8 is the number of atoms in the unit cell. Because binding energies are negligible, the Compton frequency of the sphere is $\omega_M = [m(Si)/m({}^{133}Cs)]X_{at} \omega_{Cs}$,
given by the measured Compton frequency of given by the measured Compton frequency of

Fig. 4. Compton clock performance. (A) Frequency minus 10 MHz versus time plotted over ~6 hours. (B) Histogram of data (bin size = 0.025 Hz). (C) Root Allan variance (RAV) of the data in (A). It is <10⁻⁸/(τ/1000 s)^{1/2} between integration times τ of ~100 s and ~ 1 hour. The slope between τ = 10 and 100 s is
an artifact of the 80-s undate cycle of the experiment an artifact of the 80-s update cycle of the experiment.

cesium atoms ω_{Cs} . This yields its mass as $M =$ $\omega_M \hbar/c^2$. The ratio $m(Si)/m(^{133}Cs)$ is between
the effective molar mass of the sphere's matethe effective molar mass of the sphere's material and cesium-133. According to (25), present data yields the spheres' mass with an overall accuracy of 30 ppb so that they would constitute the most accurately calibrated macroscopic masses under the proposed CGPM-2011 redefinition—a testament to the precision achieved in constructing Avogadro spheres.

Although any method for measuring microscopic mass can be used, the Compton clock offers a transparent connection between the second and a microscopic mass on the basis of simple physical principles without requiring auxiliary measurements. It directly realizes a long-standing proposal to measure mass in terms of the Compton frequency (6). The method outlined here offers a different set of systematic effects as compared with Watt balances (23, 24), thus serving as an important test of the overall consistency of the laws of physics and experimental methods. It is based on a body's inertial rather than gravitational mass. In the context of the present SI, the combination of the Avogadro project and the Compton clock serves as a measurement of the Planck constant.

Looking forward, the rapidly developing field of optomechanics (26) might enable measurements of the single-photon recoil energy of a nanomechanical mirror. This could result in a clock referenced to the mass of a mesoscopic object, or a mesoscopic mass standard. Because Bragg diffraction of electrons has already been demonstrated (27), a clock using elementary particles or even antiparticles is possible. Such clocks would be useful for testing CPT symmetry or the Einstein Equivalence Principle for antimatter.

We have demonstrated a clock stabilized to the rest mass of a particle. It highlights the intimate connection between frequency and mass. It proves that massive particles can serve as a frequency reference without requiring their mass to be converted to energy as an explicit illustration of a key principle of quantum mechanics. Furthermore, we have shown that a single massive particle is sufficient to measure time.

References and Notes

- 1. R. Penrose, Cycles of Time: An Extraordinary View of The Universe (Knopf, New York, 2011), sec. 2.3.
- 2. L. De Broglie, thesis, University of Paris, Paris, France (1924).
- 3. H. Müller, A. Peters, S. Chu, Nature 463, 926 (2010).
- 4. M. A. Hohensee, S. Chu, A. Peters, H. Müller, Phys. Rev. Lett. 106, 151102 (2011).
- 5. P. Wolf et al., Class. Quantum Gravity 28, 145017 (2011).
- 6. J. W. G. Wignall, Phys. Rev. Lett. 68, 5 (1992).
- 7. Ch. J. Bordé, Eur. Phys. J. Spec. Top. 163, 315 (2008).
- 8. M. A. Hohensee, B. Estey, P. Hamilton, A. Zeilinger, H. Müller, Phys. Rev. Lett. 108, 230404 (2012). 9. L. Essen, J. V. L. Parry, Nature 176, 280 (1955).
- 10. T. E. Parker, Metrologia 47, 1 (2010).
- 11. A. D. Ludlow et al., Science 319, 1805 (2008).
- 12. C. W. Chou, D. B. Hume, J. C. J. Koelemeij, D. J. Wineland,
- T. Rosenband, Phys. Rev. Lett. 104, 070802 (2010).
- 13. A. D. Cronin, J. Schmiedmayer, D. E. Pritchard, Rev. Mod. Phys. 81, 1051 (2009).
- 14. H. Müller, S.-W. Chiow, Q. Long, S. Herrmann, S. Chu, Phys. Rev. Lett. 100, 180405 (2008).
- 15. S.-Y. Lan, P.-C. Kuan, B. Estey, P. Haslinger, H. Müller, Phys. Rev. Lett. 108, 090402 (2012). 16. Materials and methods are available as supplementary
- materials on Science Online. 17. R. Bouchendira, P. Cladé, S. Guellati-Khélifa,
- F. Nez, F. Biraben, Phys. Rev. Lett. 106, 080801 (2011).
- 18. P. J. Mohr, B. N. Taylor, D. B. Newell, The 2010 CODATA Recommended Values of the Fundamental Physical Constants (Web Version 6.0); http://physics.nist. gov/constants.
- 19. S. Dimopoulos, P. W. Graham, J. M. Hogan, M. A. Kasevich, Phys. Rev. Lett. 98, 111102 (2007).
- 20. "Resolution 1: On the possible future revision of the International System of Units, the SI," General Conference on Weights and Measures, Sèvres, France, 17 to 21 October 2011.
- 21. D. Hanneke, S. Fogwell, G. Gabrielse, Phys. Rev. Lett. 100, 120801 (2008).
- 22. T. Aoyama, M. Hayakawa, T. Kinoshita, M. Nio, Phys. Rev. Lett. 109, 111807 (2012).
- 23. I. A. Robinson, B. P. Kibble, Metrologia 44, 427 (2007).
- 24. R. L. Steiner, E. R. Williams, R. Liu, D. B. Newell, IEEE Trans. Instrum. Meas. 56, 592 (2007).
- 25. B. Andreas et al., Phys. Rev. Lett. **106**, 030801 (2011).
- 26. T. J. Kippenberg, K. J. Vahala, Science 321, 1172 (2008). 27. D. L. Freimund, H. Batelaan, Phys. Rev. Lett. 89, 283602 (2002).

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Supplementary Materials

www.sciencemag.org/cgi/content/full/science.1230767/DC1 Materials and Methods Supplementary Text Figs. S1 and S2 Table S1 References (28–32)

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**Nanoscale Nuclear Magnetic
Resonance with a Nitrogen-Vacancy Spin Sensor**

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Extension of nuclear magnetic resonance (NMR) to nanoscale samples has been a longstanding challenge because of the insensitivity of conventional detection methods. We demonstrated the use of an individual, near-surface nitrogen-vacancy (NV) center in diamond as a sensor to detect proton NMR in an organic sample located external to the diamond. Using a combination of electron spin echoes and proton spin manipulation, we showed that the NV center senses the nanotesla field fluctuations from the protons, enabling both time-domain and spectroscopic NMR measurements on the nanometer scale.

oth nuclear magnetic resonance (NMR) spectroscopy and magnetic resonance imaging (MRI) have become indispensable tools in many diverse fields of research, including analytical chemistry, materials science, structural biology, neuroscience, and medicine (1). The one major deficiency of NMR is the low sensitivity of the conventional coil-based induction method of detection, which prevents its application to samples at the nanometer scale (2). Much improved detection sensitivity has been achieved with magnetic resonance force microscopy, which is based on detecting weak magnetic forces in

the presence of a strong field gradient and has demonstrated nanometer-scale NMR imaging at cryogenic temperatures (3). Here, we present an alternative nanoscale detection method that works in the absence of a magnetic field gradient, thus preserving spectroscopic information, and is operable over a wide range of temperatures, including room temperature. A single near-surface nitrogen-vacancy (NV) center in diamond is used as an atomic-size sensor to detect weak magnetic fields originating from nuclear spins external to the diamond. In an initial demonstration, we detected randomly polarized hydrogen nuclei (protons) in an organic polymer. Both time domain and spectroscopic information were obtained by appropriately manipulating the protons so as to affect the precession phase of the highly coherent NV electron spin. The results suggest that NV-based NMR detection may provide a path toward three-dimensional nanoscale magnetic resonance imaging (nanoMRI) under ambient conditions $(4, 5)$.

NV centers are proving to be particularly useful for both quantum information processing and nanoscale magnetic sensing $(6-11)$. The negatively charged center has a spin state with a long coherence time, especially in isotopically purified crystals (12, 13), and an electronic-level structure

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